

Let's start by discussing quantum noise.

Similar to the classical spectral density, we define a quantum spectral density:

$$S_{xx}(\omega) = \int_{-\infty}^{\infty} dt e^{i\omega t} \langle \hat{x}(t) \hat{x}(0) \rangle$$

$\hat{x}$  is a quantum operator

$\langle \rangle$  is a quantum statistical average using a density matrix

$$\rightarrow \frac{d\hat{A}}{dt} = \frac{i}{\hbar} [\hat{H}, \hat{A}] + \frac{\partial \hat{A}}{\partial t}$$

Solutions to the Heisenberg equations of motion

for a harmonic osc. give:

$$\hat{x}(t) = \hat{x}(0) \cos(\omega_0 t) + \hat{p}(0) \frac{1}{m\omega_0} \sin(\omega_0 t)$$

$$\hat{p}(t) = \hat{p}(0) \cos(\omega_0 t) - \hat{x}(0) m\omega_0 \sin(\omega_0 t)$$

$$\therefore \langle \hat{x}(t) \hat{x}(0) \rangle = \langle \hat{x}(0) \hat{x}(0) \rangle \cos(\omega_0 t) + \langle \hat{p}(0) \hat{x}(0) \rangle \frac{1}{m\omega_0} \sin(\omega_0 t)$$

Recall, however, that:  $[\hat{x}(0), \hat{p}(0)] = i\hbar$

$$\hat{x} = x_{ZPF} (\hat{a}^+ + \hat{a})$$

$$\hat{p} = \frac{i\hbar}{2x_{ZPF}} (\hat{a}^+ - \hat{a})$$

$$\text{w/ } [\hat{a}, \hat{a}^+] = 1$$

$$\hat{N} = \hat{a}^+ \hat{a}$$

$$\text{w/ } x_{ZPF}^2 = \langle 0 | \hat{x}^2 | 0 \rangle = \frac{\hbar}{2m\omega_0}$$

For thermal equilibrium:

$$\langle \hat{x}(0) \hat{p}(0) \rangle = +i \frac{\hbar}{2}$$

$$\langle \hat{p}(0) \hat{x}(0) \rangle = -i \frac{\hbar}{2}$$

So, the correlator is:

$$\langle \hat{x}(t) \hat{x}(0) \rangle = \langle \hat{x}(0) \hat{x}(0) \rangle \frac{e^{i\omega_0 t} + e^{-i\omega_0 t}}{2} + \langle \hat{p}(0) \hat{x}(0) \rangle \frac{1}{m\omega_0} \frac{e^{i\omega_0 t} - e^{-i\omega_0 t}}{2i}$$

$$\hat{x} \hat{x} = x_{zpf}^2 (\hat{a}^\dagger \hat{a}^\dagger + \hat{a}^\dagger \hat{a} + \hat{a} \hat{a}^\dagger + \hat{a} \hat{a})$$

$$\hat{p} \hat{x} = \frac{i\hbar}{2} (\hat{a}^\dagger \hat{a}^\dagger + \hat{a}^\dagger \hat{a} - \hat{a} \hat{a}^\dagger - \hat{a} \hat{a})$$

$$\langle \hat{x}(t) \hat{x}(0) \rangle = x_{zpf}^2 \left[ \langle (2\hat{n} + 1) \rangle \frac{e^{i\omega_0 t} + e^{-i\omega_0 t}}{2} - \langle (1) \rangle \frac{e^{i\omega_0 t} - e^{-i\omega_0 t}}{2} \right]$$

$$\langle \hat{x}(t) \hat{x}(0) \rangle = x_{zpf}^2 \left[ \langle \hat{n} \rangle e^{i\omega_0 t} + \langle \hat{n} + 1 \rangle e^{-i\omega_0 t} \right] \leftarrow$$

∴ The spectral density is therefore:

$$(1) \quad S_{xx}(\omega) = 2\pi x_{zpf}^2 \left\{ \langle \hat{n} \rangle \delta(\omega + \omega_0) + \langle \hat{n} + 1 \rangle \delta(\omega - \omega_0) \right\}$$

↑  
stimulated emission  
into the osc.

(absorption)

↑  
emission  
by the osc

(emission)

In the high temperature limit:

$$k_B T \gg \hbar \Omega$$

$$\langle \hat{n} \rangle \sim \langle \hat{n} + 1 \rangle \sim \frac{k_B T}{\hbar \omega_0}$$

$$S_{xx}(\omega) = \pi \frac{k_B T}{m \omega_0^2} \left\{ \delta(\omega + \omega_0) + \delta(\omega - \omega_0) \right\}$$

Let's take a general cavity  
(optical or RF) coupled to the  
position of an oscillator.

Changes in the position of the oscillator lead  
to changes in the phase of the cavity's  
reflected carrier signal. The changing phase  
shift can be converted to a changing  
intensity (by interferometry).

Applies to microwave cavities, optical  
cavities, etc.

cavity → uses wave nature of light  
for the interference signal

• suffers from particle nature  
of light in the noise.

A coherent photon state contains a Poisson distribution  
of the number of photons:

$$(\Delta N)^2 = \bar{N}$$

↑
↑  
 fluct.                      mean

$$(\Delta \theta)^2 = \frac{1}{4\bar{N}}$$

↑  
 phase  
 uncertainty

⇒  $\Delta N \Delta \theta = \frac{1}{2}$

Number-phase uncertainty  
relation.

In spectral densities this gives:

$S_{\sin S_{00}} = \frac{1}{4}$  spectral density of phase fluctuations  
 $\sqrt{S_{\sin S_{00}}} = \frac{1}{2}$  spectral density of photon flux

Take the case of light reflecting off a mirror. The beam will have a phase shift  $2kx$  if the mirror moves  $x$ . The phase uncertainty leads to a position imprecision:

$$S_{xx}^I = \frac{1}{4k^2} S_{00}$$

The back action imparted by a photon hitting the mirror is  $2\hbar k$ . So the photon shot noise corresponds to a random back action force:

$$S_{FF} = 4\hbar^2 k^2 S_{\sin}$$

Together:

$$S_{FF} S_{xx}^I = \hbar^2 S_{\sin} S_{00} = \frac{\hbar^2}{4} \quad (2)$$

$$\sqrt{S_{FF} S_{xx}^I} = \frac{\hbar}{2}$$

Quantum limit on noise of detector

Let's apply all of this to our mechanical oscillator in a resonant cavity.

Standard Quantum Limit for Position Measurement

(weak measurement - many cycles of oscillation before information is acquired)

$$\hat{x}(t) = \hat{X} \cos(\omega_0 t) + \hat{Y} \sin(\omega_0 t)$$

$$[\hat{x}, \hat{y}] = [\hat{x}(0), \frac{1}{m\omega_0} \hat{p}(0)] = \frac{1}{m\omega_0} [\hat{x}(0), \hat{p}(0)] = \frac{i\hbar}{m\omega_0}$$

$$[\hat{x}, \hat{y}] = 2i X_{ZPF}^2$$

B/c this is a weak measurement, we are trying to measure two incompatible observables simultaneously.

Spectral Density of harmonic oscillator w/ resonant frequency  $\omega_0$  & damping  $\gamma_0$  & equation of motion

$$m\ddot{x} + m\gamma_0\dot{x} + m\omega_0^2 x = F$$

$$x(t) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} x(\omega) e^{-i\omega t} d\omega$$

is

$$S_{xx}(\omega) = X_{ZPF}^2 \left\{ \langle N \rangle \frac{\gamma_0}{(\omega + \omega_0)^2 + (\frac{\gamma_0}{2})^2} + \langle N+1 \rangle \frac{\gamma_0}{(\omega - \omega_0)^2 + (\frac{\gamma_0}{2})^2} \right\}$$

where we have referred back to (1) on page 2 and converted  $\delta$ -functions into Lorentzians.

For the classical limit ( $k_B T \gg \hbar\omega_0$ ):

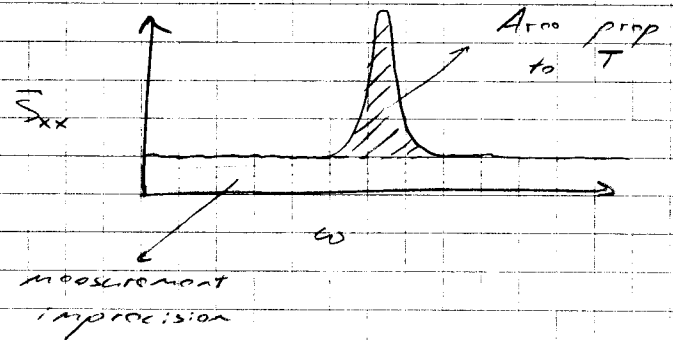
$$\bar{S}_{xx}(\omega) = \frac{1}{2} [S_{xx}(\omega) + S_{xx}(-\omega)]$$

$$\approx \frac{k_B T}{2m\omega_0^2} \frac{\gamma_0}{(\omega - \omega_0)^2 + (\frac{\gamma_0}{2})^2}$$

$$\langle x^2 \rangle = \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} \bar{S}_{xx}(\omega) = \frac{k_B T}{m\omega_0^2}$$

$$\therefore m\omega_0^2 \langle x^2 \rangle = k_B T$$

$$\hookrightarrow \frac{1}{2} m\omega_0^2 \langle x^2 \rangle = \frac{1}{2} k_B T \quad \checkmark \text{ Equipartition}$$



At zero temperature, we have:

$$\bar{S}_{xx}^0 = X_{ZPF}^2 \frac{\delta\omega/2}{(\omega - \omega_0)^2 + (\delta\omega/2)^2}$$

One might expect to measure this, but this neglects the effect of back-action.

A displacement  $S_x(\omega)$  will be caused by a force  $F(\omega)$  through the mechanical susceptibility  $\chi_{xx}(\omega)$ .

$$S_x(\omega) = \chi_{xx}(\omega) F(\omega)$$

From:

$$m\ddot{x} + m\gamma_0\dot{x} + m\omega_0^2 x = F$$

$$\text{w/ } \chi_{xx}(\omega) = \frac{1}{m} \frac{1}{\omega_0^2 - \omega^2 - i\gamma_0\omega}$$

So, extra displacement will be caused by the back-action of the detector.

Let's look at the resonant freq. of the oscillator,  $\omega = \omega_0$ :

$$\bar{S}_{xx, tot}(\omega_0) = \bar{S}_{xx}^0(\omega_0) + \frac{1}{2} |K_{xx}(\omega_0)|^2 (S_{FF}(\omega_0) + S_{FF}(-\omega_0)) + \frac{1}{2} (S_{xx}^I(\omega_0) + S_{xx}^I(-\omega_0))$$

back-action

$$\bar{S}_{xx, tot}(\omega_0) = \bar{S}_{xx}^0(\omega_0) + \bar{S}_{xx, add}(\omega_0)$$

$\bar{S}_{xx}^0$  is the zero-point motion:

measurement imprecision

$$\bar{S}_{xx}^0(\omega_0) = \frac{2 \times 20^2}{80}$$

Let's assume the best possible detector — a quantum limited detector. See (2) from (4):

$$S_{FF} S_{xx}^I = \frac{\hbar^2}{4}$$

$$i. \bar{S}_{xx, add}(\omega_0) = |K_{xx}(\omega_0)|^2 S_{FF} + \frac{\hbar^2}{4} \frac{1}{S_{FF}}$$

least additional noise created when:

$$S_{FF, opt} = \frac{\hbar}{2 |K_{xx}(\omega_0)|} = \frac{\hbar}{2 m \omega_0 \gamma_0}$$

for large damping, less suscept. to back-action

At optimal point:

$$\bar{S}_{xx, add}(\omega_0) = \frac{\hbar}{m \omega_0 \gamma_0} = \bar{S}_{xx}^0(\omega_0)$$

The added position noise is exactly equal to the noise power of the zero-point fluctuations.

This limit has nothing to do w/ temperature; this is purely a quantum effect. If you have a perfect detector ( $S_{FF} S_{xx}^I$  is minimum), this will be the noise limit.

$$\therefore \overline{S_{xx,tot}}(\omega_0) = 2 \overline{S_{xx}^0}(\omega_0) \quad \Leftarrow$$

↑  
total intracord position noise

- 1/2 noise from the oscillator
- 1/2 noise from the detector

If we consider all figures, not just the resonance  $\omega = \omega_0$ , then:

$$\begin{aligned} \overline{S_{xx,tot}}(\omega) &= x_{2PF}^2 \frac{\delta_0/2}{(\omega - \omega_0)^2 + (\delta_0/2)^2} + \frac{\hbar}{2} \left( \frac{|X_{xx}(\omega)|^2}{|X_{xx}(\omega_0)|} + |X_{xx}(\omega_0)| \right) \\ &\approx \frac{x_{2PF}^2}{\delta_0} \left\{ 1 + 3 \frac{(\delta_0/2)^2}{(\omega - \omega_0)^2 + (\delta_0/2)^2} \right\} \end{aligned}$$



$$\frac{d\hat{A}}{dt} = \frac{i}{\hbar} [\hat{H}, \hat{A}] + \frac{\partial \hat{A}}{\partial t}$$

$$\hat{H} = \frac{\hat{p}^2}{2m} + \frac{m\omega_0^2 \hat{x}^2}{2}$$

$$\frac{d}{dt} \hat{x} = \frac{i}{\hbar} [\hat{H}, \hat{x}] = \frac{\hat{p}}{m}$$

$$\frac{d}{dt} \hat{p} = \frac{i}{\hbar} [\hat{H}, \hat{p}] = -m\omega_0^2 \hat{x}$$

$$\dot{\hat{p}}(0) = -m\omega_0^2 \hat{x}(0)$$

$$\dot{\hat{x}}(0) = \frac{p(0)}{m}$$

$$\hat{x}(t) = \hat{x}(0) \cos(\omega_0 t) + \frac{\hat{p}(0)}{m\omega_0} \sin(\omega_0 t)$$

$$\hat{p}(t) = \hat{p}(0) \cos(\omega_0 t) - m\omega_0 \hat{x}(0) \sin(\omega_0 t)$$